



Noble Element Simulation Technique

for Geant4

<http://nest.physics.ucdavis.edu>

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on behalf of the entire NEST development team, of the University of California, Davis, Davis, CA, USA, Lawrence Livermore National Laboratory, Livermore, CA, USA, and the University of California, Los Angeles, CA, USA

University of Minnesota, Minneapolis, Friday 11-11-11

The People of the NEST Team

UC Davis, LLNL, and UCLA, in California

A small but passionate group of individuals who love the work

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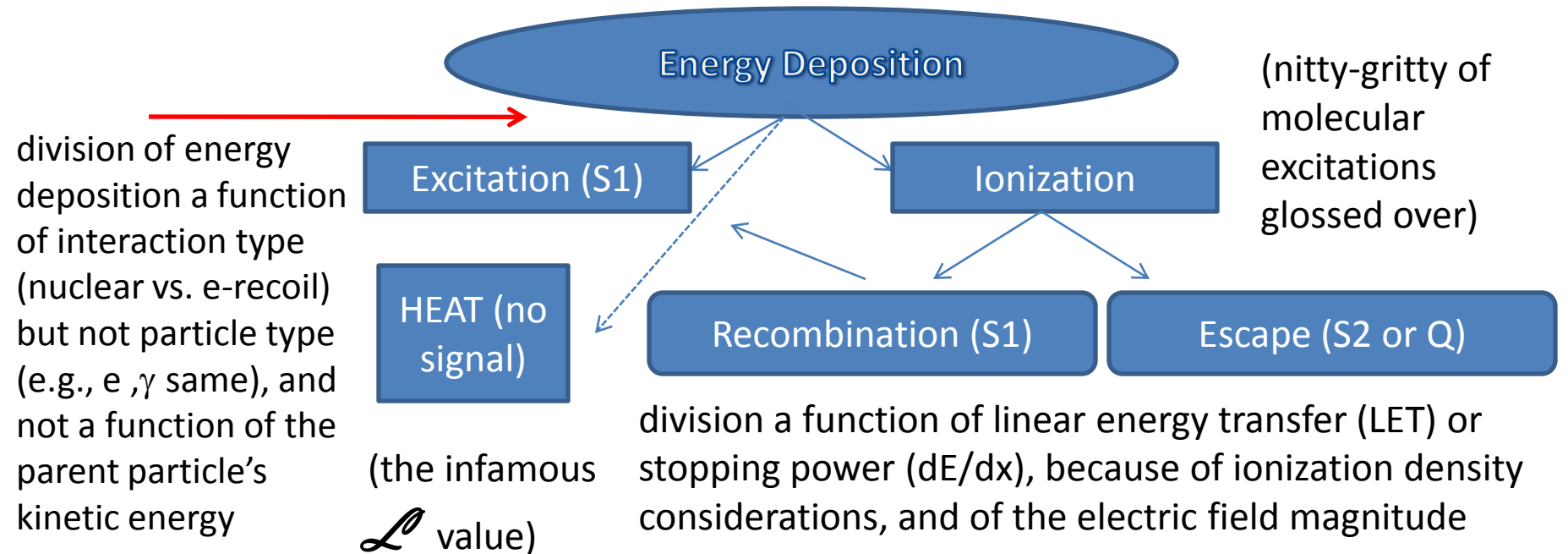
Why simulate scintillating noble elements well?

- Direct dark matter detection or calibration for it (past, present, future experiments)
 - LUX, XENON, ZEPLIN, LZ, WArP, DarkSide, ArDM, XMASS, DARWIN, MAX, Xürich, Xed, XeCube, PANDA-X, PIXeY, DEAP, CLEAN, 1- and 2-phase
- Double beta decay ($0\nu\beta\beta$, $2\nu\beta\beta$) projects too
 - EXO, NEXT (both ^{136}Xe -enriched)
- Positron Emission Tomography (PET) scans for medical applications: detect 511 keV γ 's
- Other particle detection applications, e.g., collider experiments (MEG, Olive, et al.)

The Purpose and Scope of NEST

- Create full-fledged simulation based on a physical, albeit also heuristic/quasi-empirical approach
- Comb the wealth of data for liquid and gaseous noble elements for different particles, energies, and electric fields, and then combine everything
- Aid the many dark matter, double beta decay, and other experiments which utilize this technology to be on the same or comparable page for simulations
- Bring added realism to the simple model that is present now in Geant4 for scintillation
- Explore backgrounds at low energy by expanding Geant4 physics to be more accurate when you go to a low energy regime: $O(1)$ keV and even lower
- Have to start somewhere: LXe, for sake of LUX

Basic Physics Principles



- Heat loss for nuclear recoils (Lindhard effect), while electron recoils relatively easier to deal with
- Start simple: no exotic energy loss mechanisms (like “bi-excitonic” collisions). Explains the data?

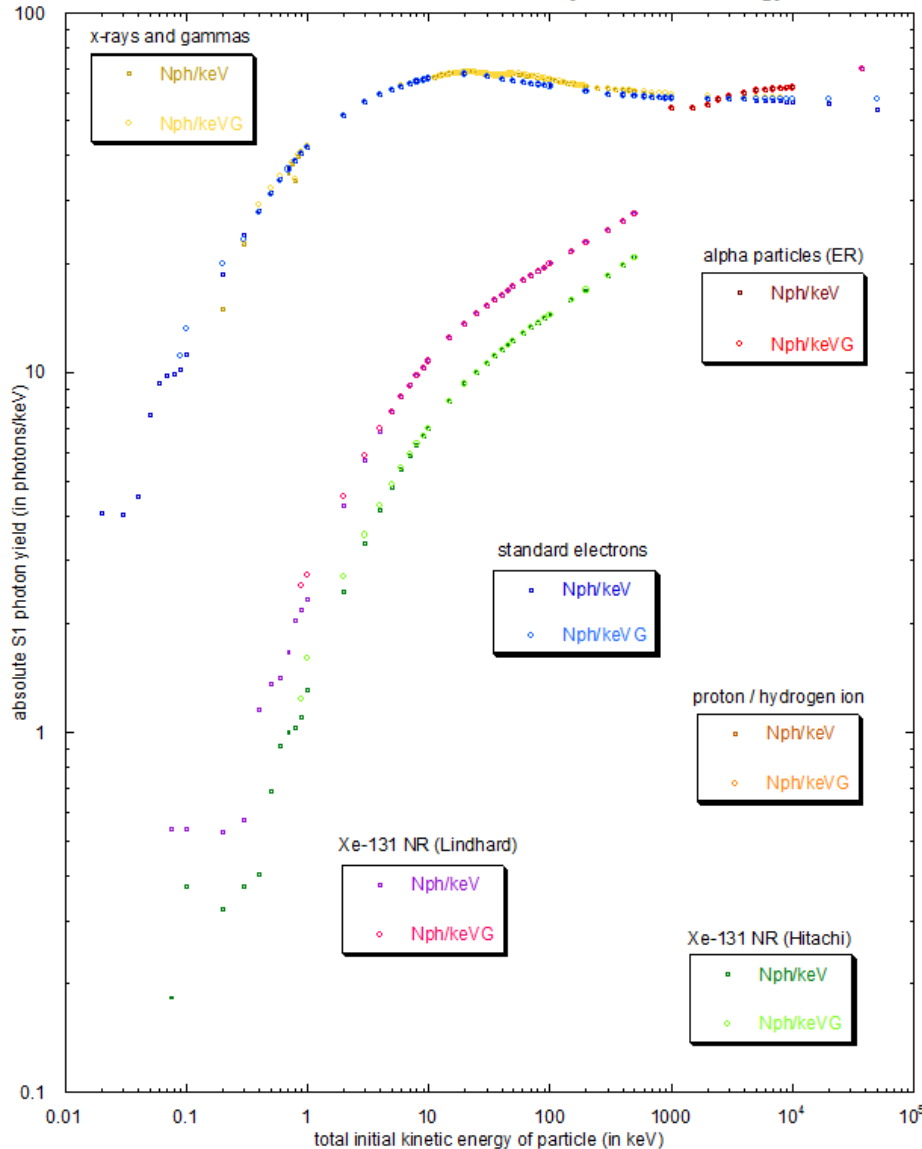
Basic Physics Principles

- Cornerstone: There is ONE work function for production of EITHER a scintillation photon or an ionization electron. All others derive from it.
- $W = 13.7 \pm 0.2 \text{ eV} \quad (N_{e^-} + N_{\gamma}) = E_{\text{dep}} / W$
C.E. Dahl, Ph.D. Thesis, Princeton University, 2009
- Two recombination models, for different E regimes
 - Thomas-Imel "box" model ($O(10)$ keV and lower)
 - Modified Birks' Law of Doke ($O(100)$ keV and up)
- Recombination probability makes for non-linear yield: 2x energy does not mean 2x light, charge
- Excellent vetting against much past data

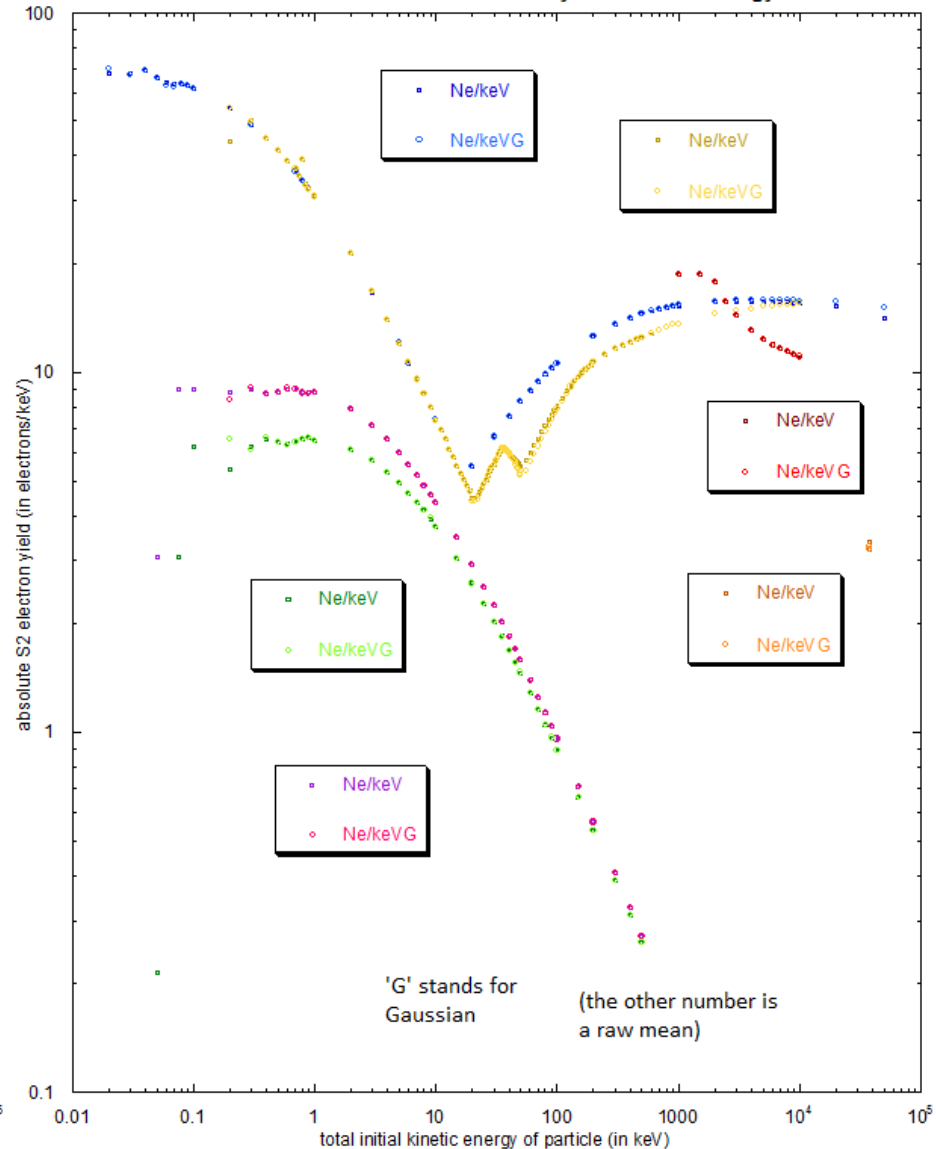
Putting it All Together to Predict the Yield

First: Let's look at zero-field scintillation yield from various particles

Zero-Field NEST Scintillation Yields by Particle and Energy



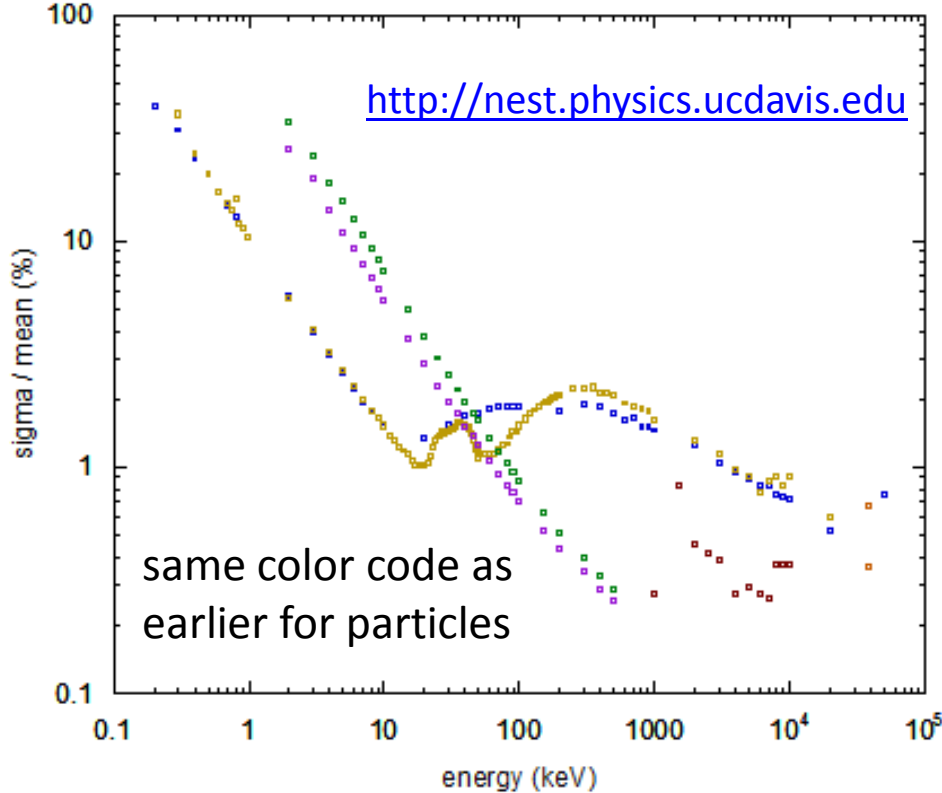
Zero-Field NEST Ionization Yields by Particle and Energy



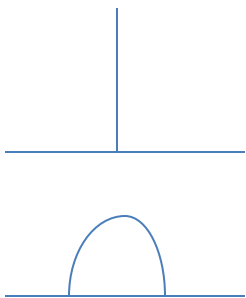
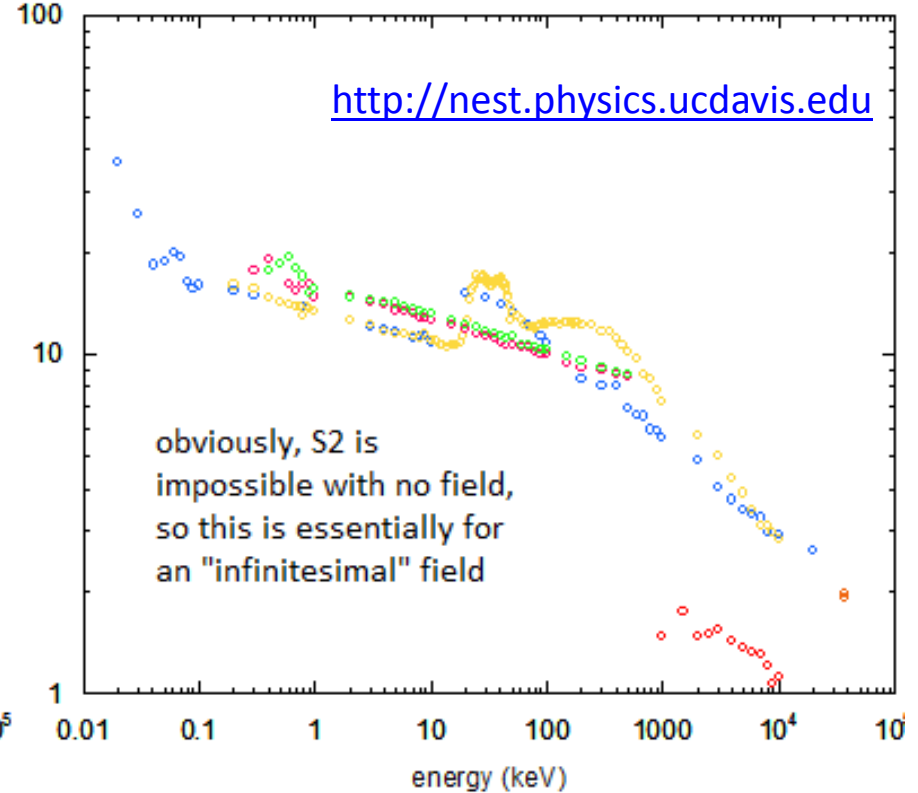
Energy Resolution

Preliminary NEST Predictions for Zero Electric Field

S1 Energy Resolution

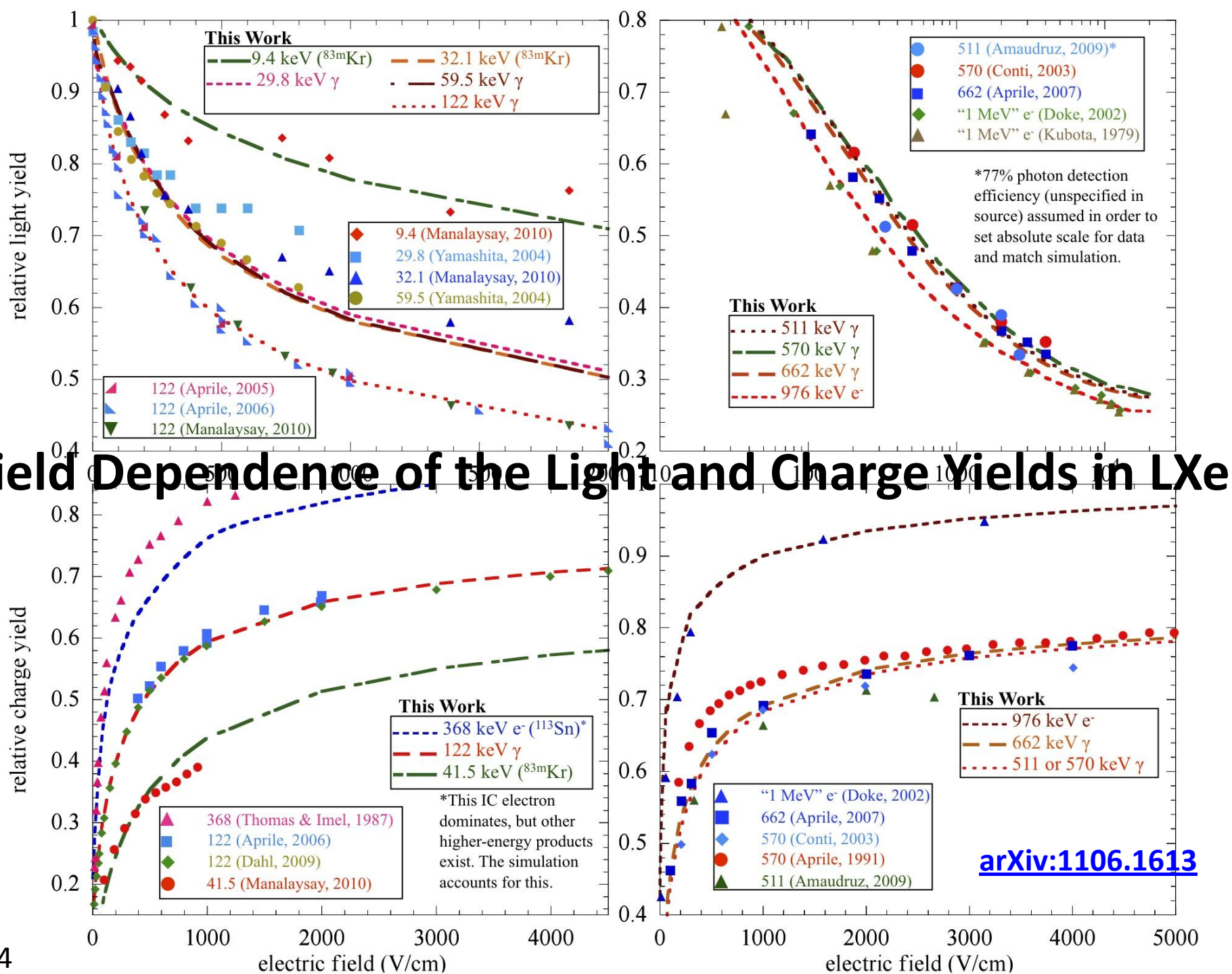


S2 Energy Resolution



The sources of INTRINSIC non-perfect resolution, at all electric fields

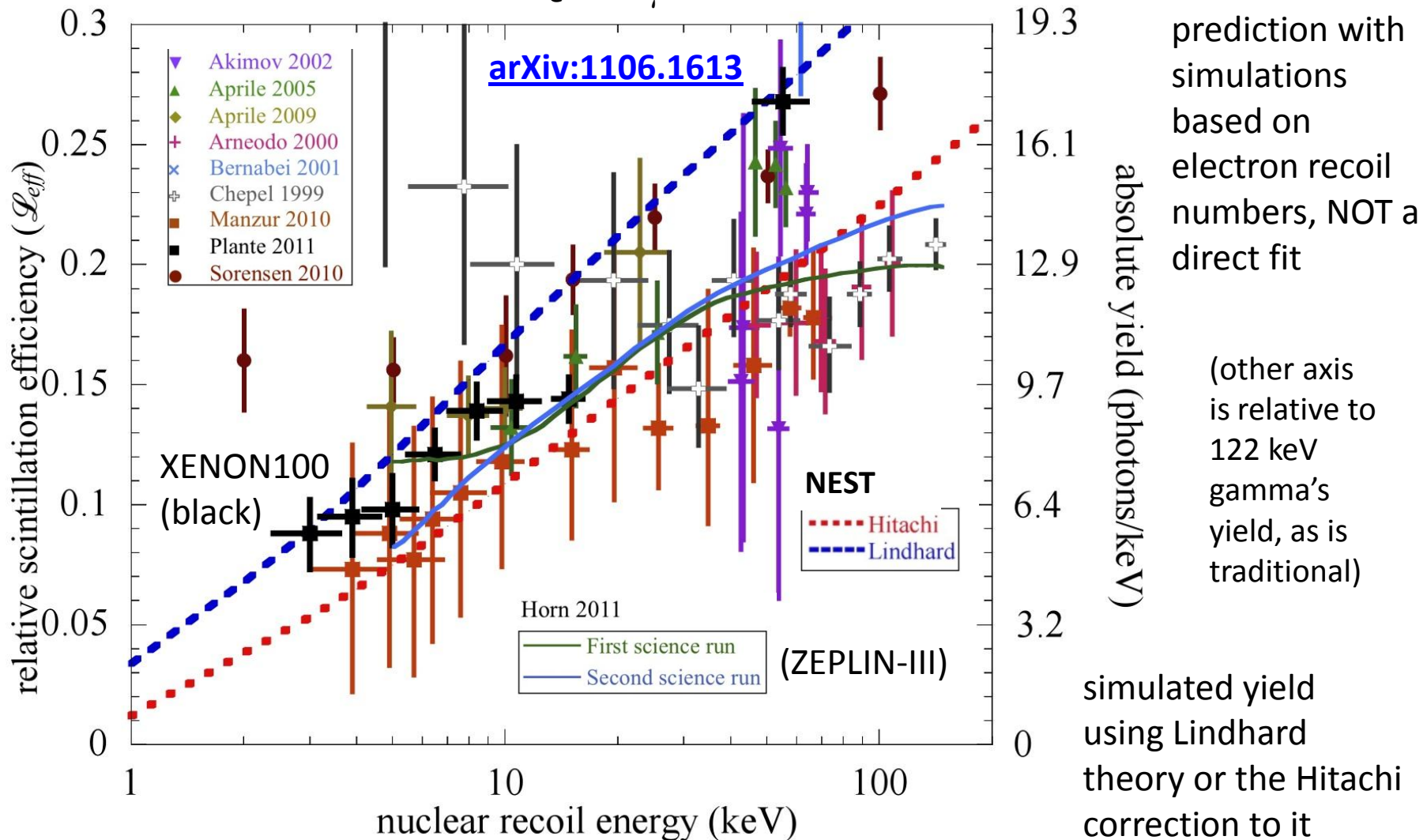
- Fano factor (extremely small effect until lower energies)
- Binomial fluctuations in the recombination probability
- Binomial fluctuations in the numbers of excitons versus ions (small)
- Particle track history, including stochastic dE/dx effects
- Dependence of recombination probability on track angle (?)



[arXiv:1106.1613](https://arxiv.org/abs/1106.1613)

The Tricky Nuclear Recoil Issue: LXe

$$(N_{e^-} + N_{\gamma}) = L(E) * E / W$$



This is likely the strongest prediction, with the simplest assumptions, ever devised!

Very Preliminary LAr Result for NR Light Yield

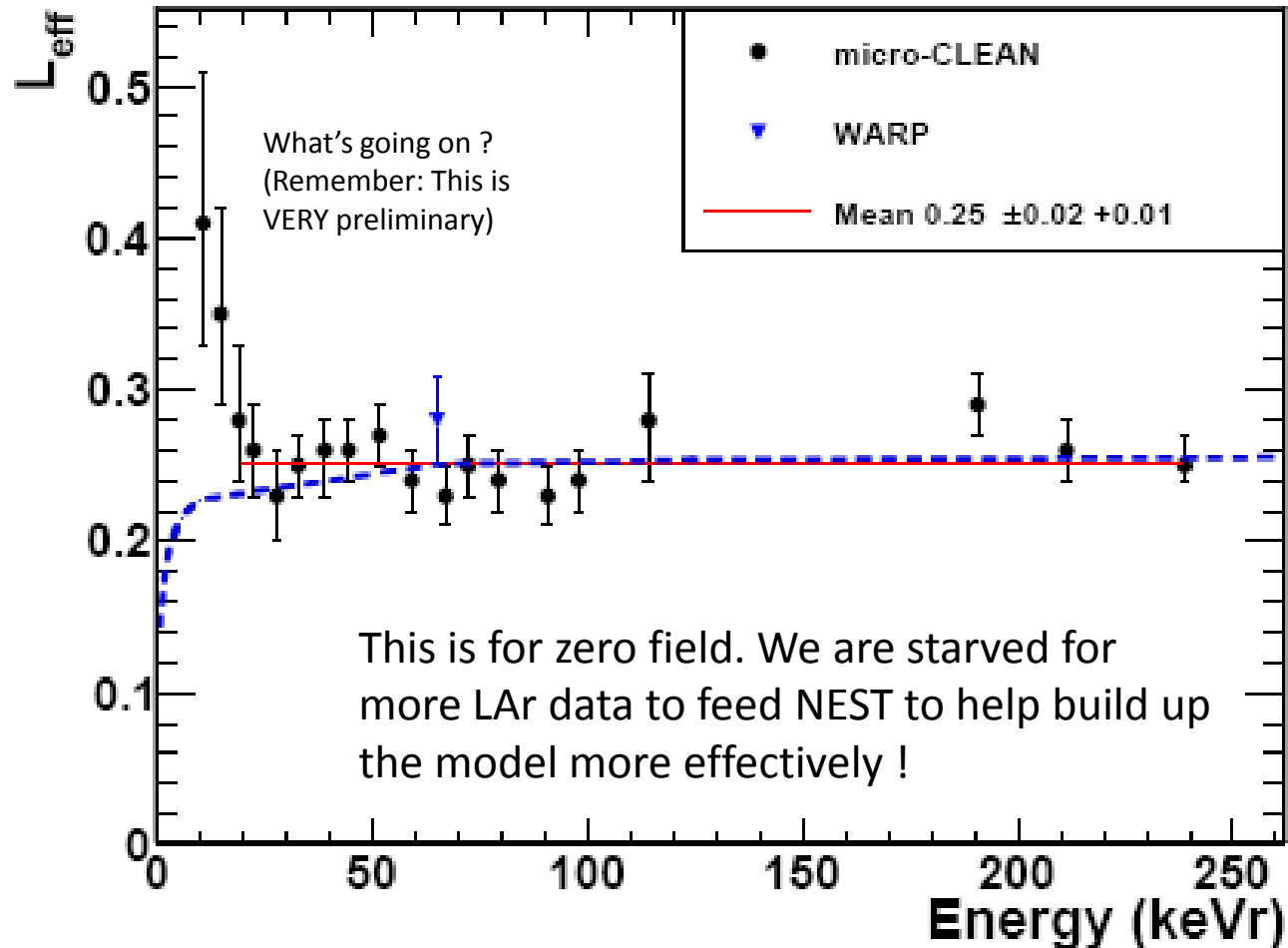
Using a flat eye-balled valued for $L(E)$
plus the Thomas-Imel box model
of recombination, with an ansatz
calculation for its one free parameter

Yield higher
than in Xe.

Why?

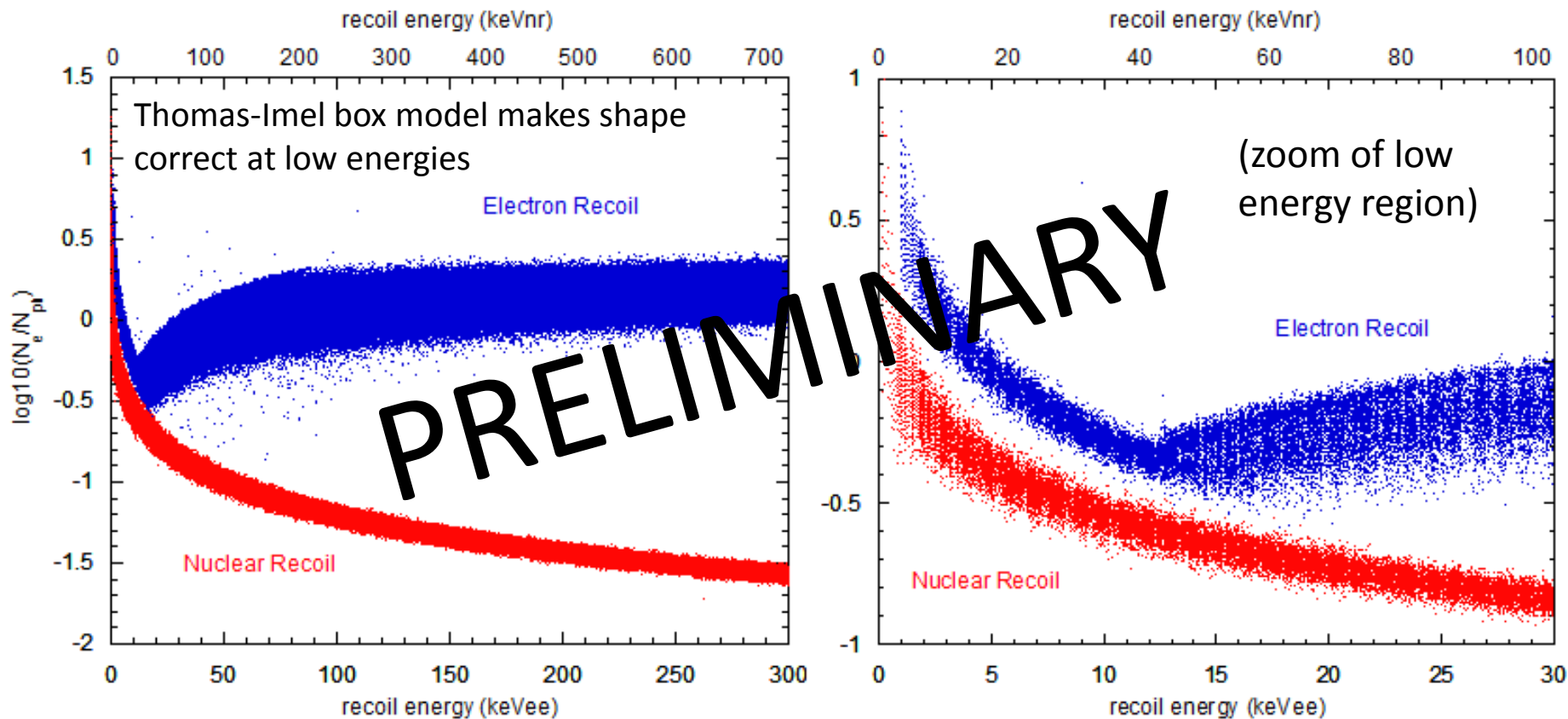
Two main
reasons:

Smaller Z
and A,
greater
initial
exciton to
ion ratio



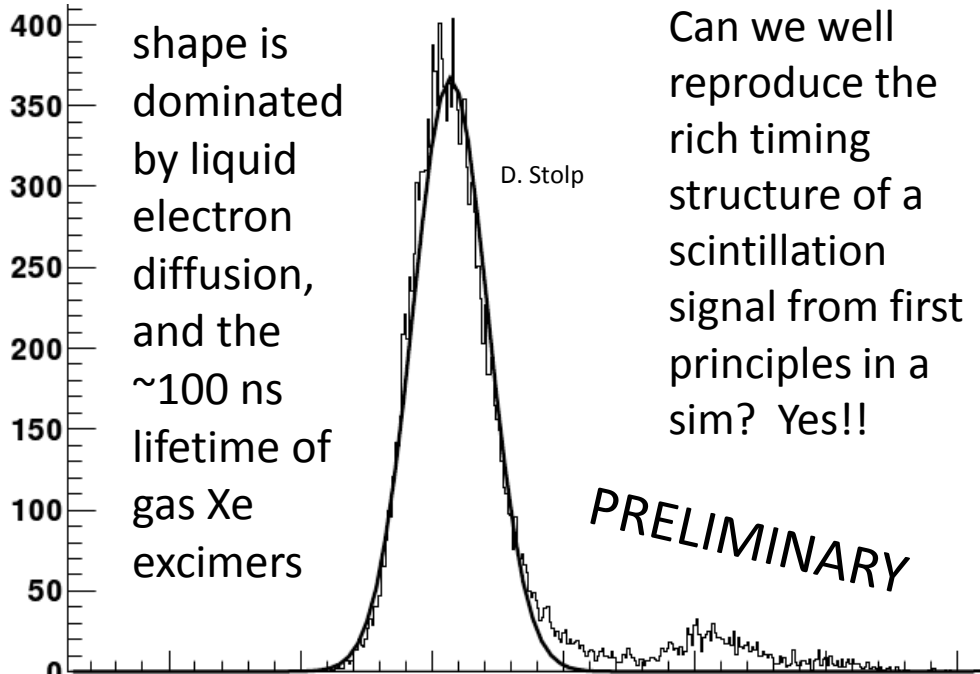
What Can NEST Do For You?

Example simulated NR, ER bands in S2/S1 space from LUX (with LUXSim - see Kazkaz' talk)



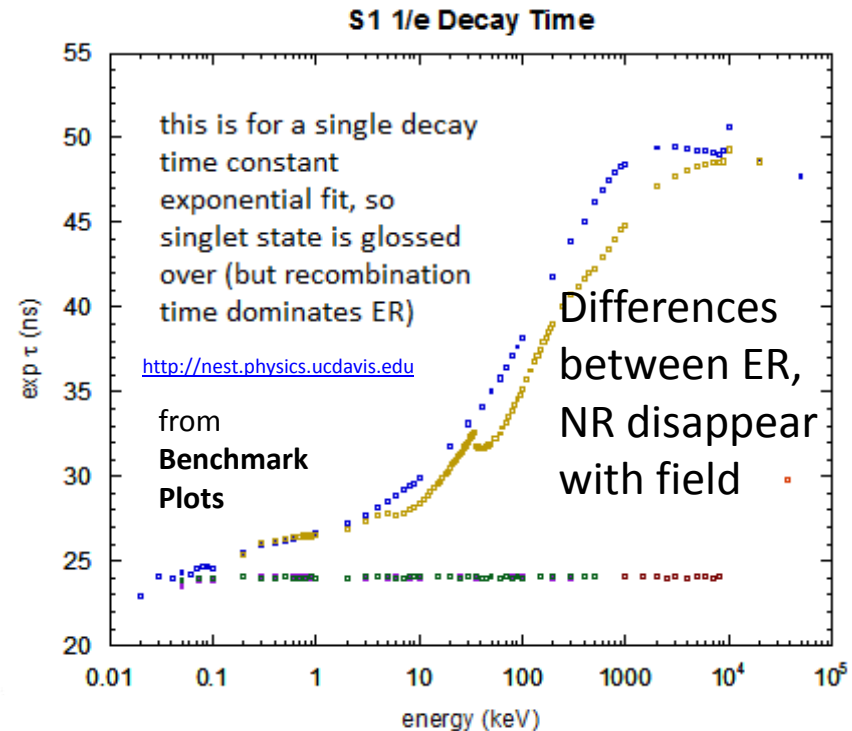
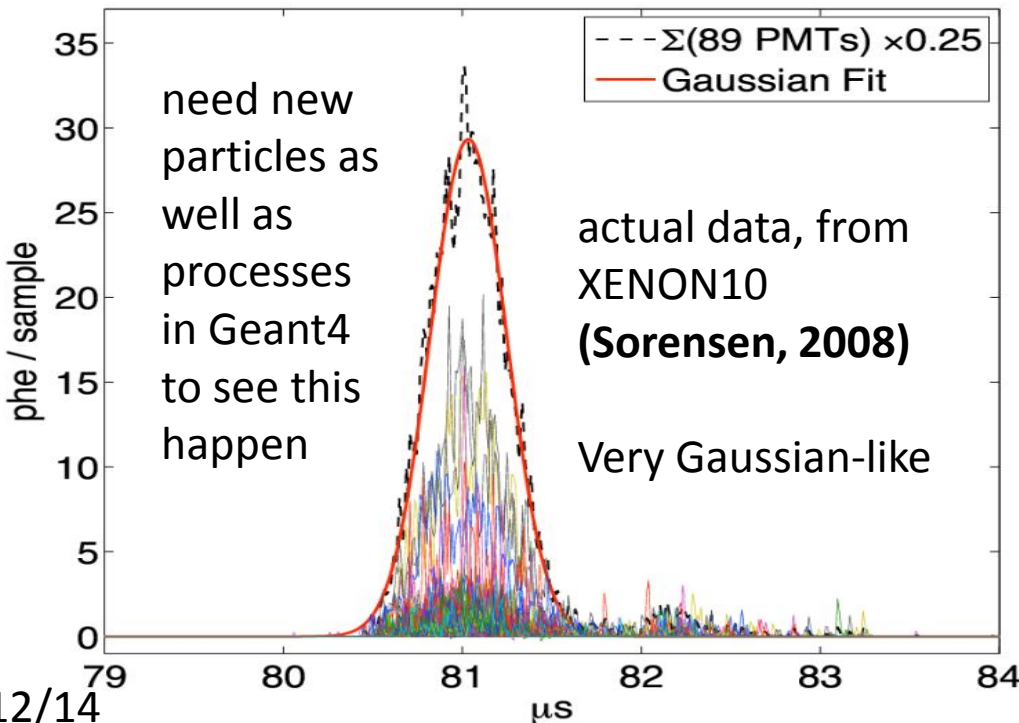
NO artificial smearing, Gaussian or otherwise, was added to NEST to yield the result depicted.

Now it has become possible, with NEST, to study/predict the discrimination power of your experiment before you even built it or calibrate, with a reliable simulation.



Understanding the Raw Pulse Shapes (S1, S2)

single, triplet lifetimes incorporated for S1, as well as the recombination time, which is varied versus ionization density and the electric field magnitude



Interfacing with GEANT

- Upgrade to **G4Scintillation** physics process, called **G4S1Light**, available for download on-line; speaking with GEANT about inclusion in upcoming version
- Another new G4 physics process: **G4S2Light** soon, utilizing a new species of electron (“thermal”)
- Resolving geometrical bugs: I want to talk to the representatives of GEANT present here today
 - Not all photons collected when reflectivities set to 100%
 - Volume borders “fuzzy” at $O(0.1 \text{ mm})$ level
- Bugs solvable, and *you can dial in a particle type and energy, set your electric field, and watch your sims give reliable results with relative ease*

Status and Future

- Fully simulating DAQ chain (pulse shaping effects)
- Detector effects: QE, PMT gain variation, etc.
- Representatives of many collaborations already members of the NEST mailing list, and downloading
- No more rules of thumb, nor extrapolations from past detectors: build your geometry and go
- Drag and drop NEST into your sim with minimal effort, and get regular updates with more features
- Next: **GXe**, **L/GAr**, Ne, He, Kr, solids – complete set someday soon? ALL noble elements...

References

- For all of the references used in this talk, please simply consult the full bibliography of

M. Szydagis et al., NEST: A Comprehensive Model
For Scintillation Yield in Liquid Xenon

2011 **JINST** 6 P10002

Also, visit us on the arxiv at:

[arxiv:1106.1613](https://arxiv.org/abs/1106.1613)

Bonus Slides

